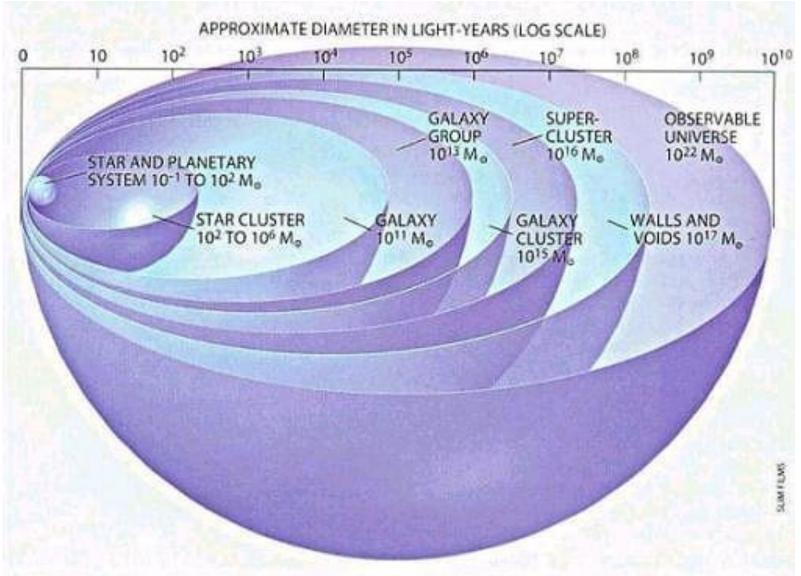
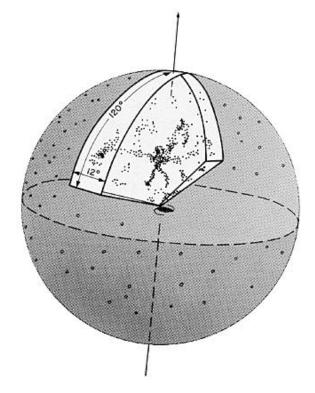
<u>Large Scale Structure</u> <u>in the Universe</u>

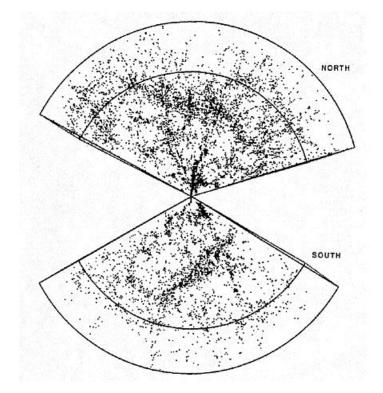
- We seem to be located at the edge of a Local Supercluster, which contains dozens of clusters and groups over a 40 Mpc region.
- Galaxies and clusters seem to congregate in "sheets", with extensive intervening "voids"
- Must reflect on how clusters of galaxies form
- Are the voids *real* voids, or could they be filled with dark matter?



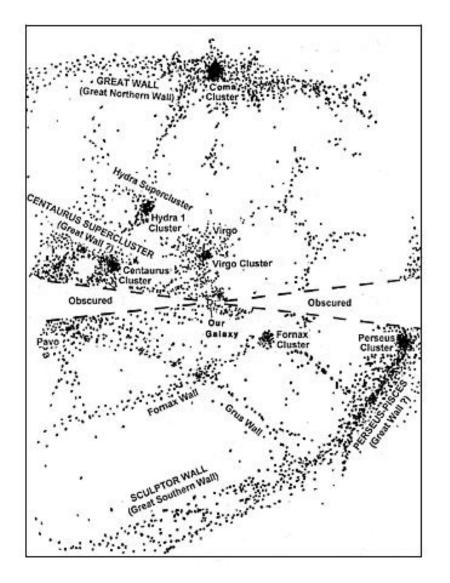
HIERARCHY OF COSMIC STRUCTURES ranges from stars and planets to the universe itself. The largest objects held together by gravity are galaxy clusters with masses up to 10^{15} times that of the sun (denoted as M_{\odot}). Although there is a higher level of organization consisting of superclusters and great walls, these patterns are not bound gravitationally. On even larger scales, the universe is featureless. Astronomers think most of these structures form from the progressive agglomeration of smaller units.

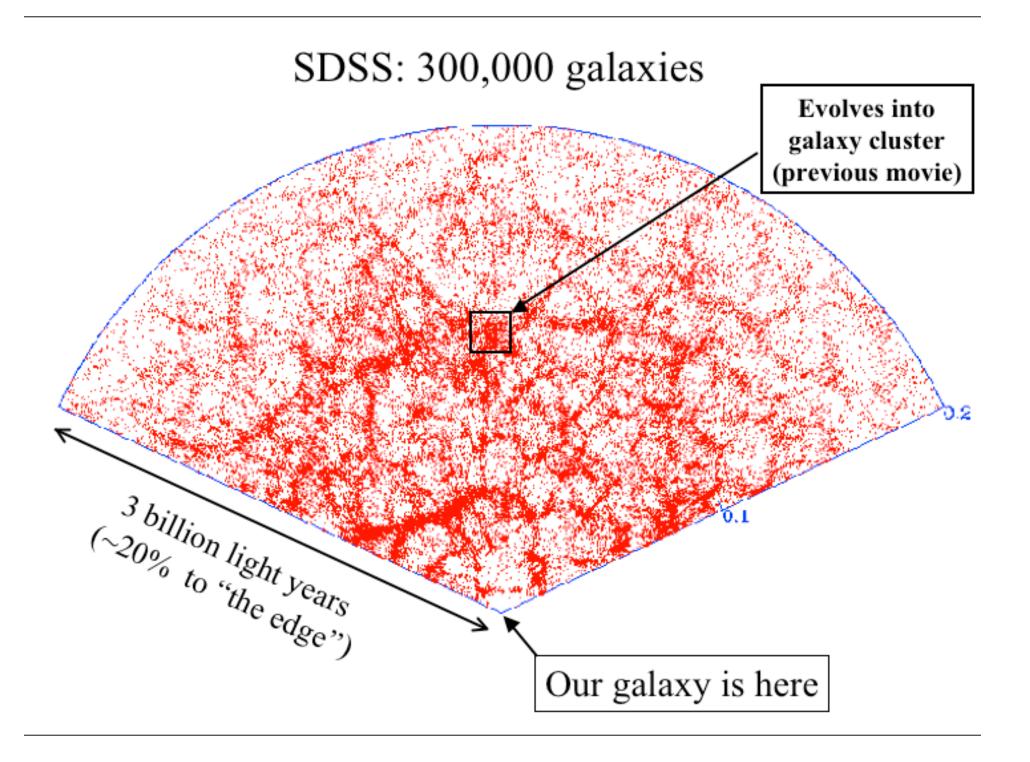
Galaxy Counting





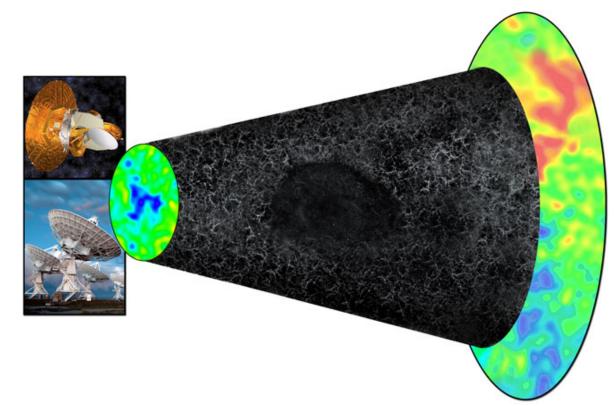
Mapping How Galaxies Are Distributed





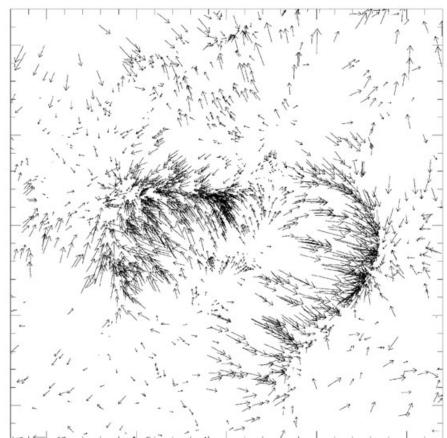
Biggest Void in Town

NRAO radio mapping has revealed a sizable region of space about 1 billion LY across that is devoid of galaxies, gas, and dark matter. It is About 6–10 billion LYs from Earth.

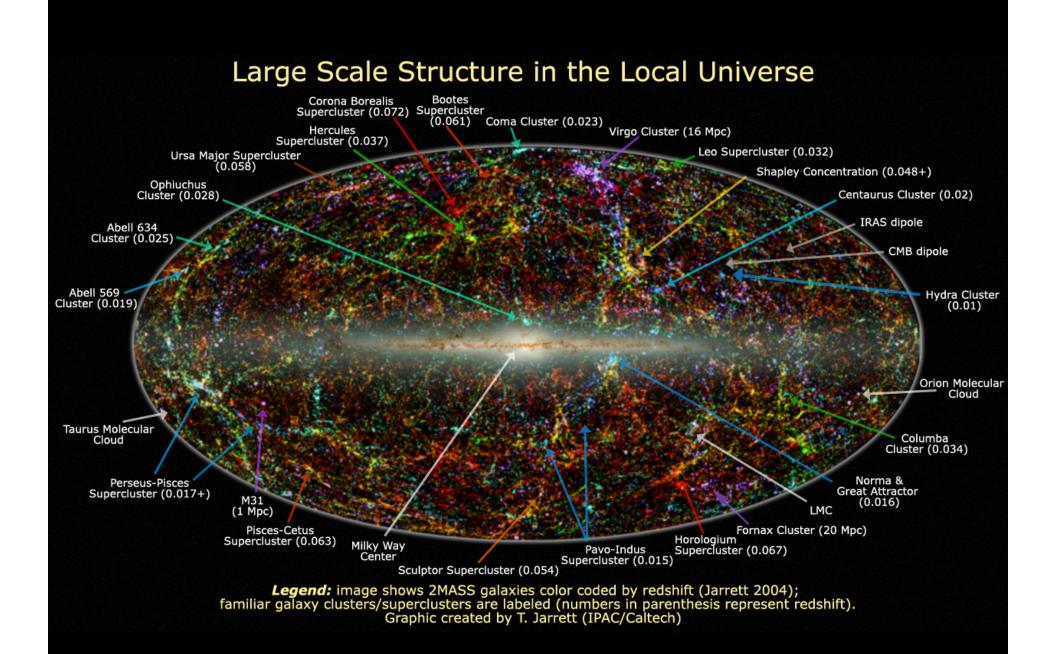


<u>Comment on "Peculiar" Velocities</u>

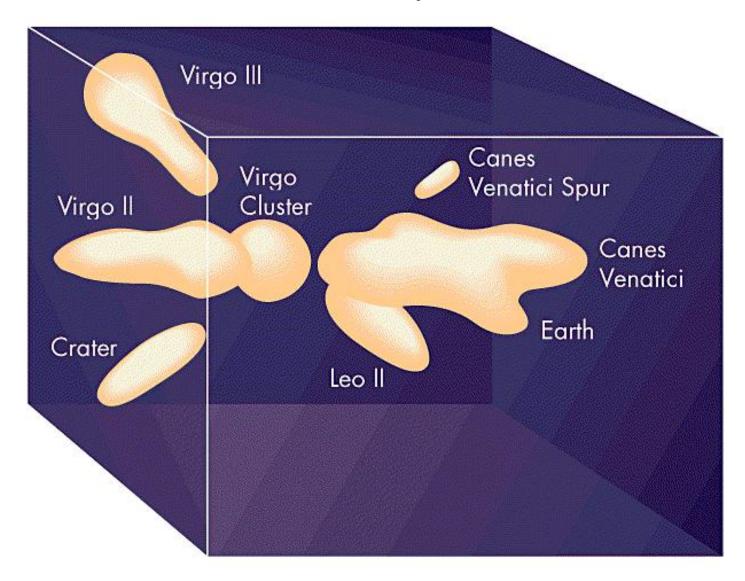
In clusters of galaxies bound by mutual gravity, the cluster participates in the Hubble flow, but the individual galaxies move in bound orbits



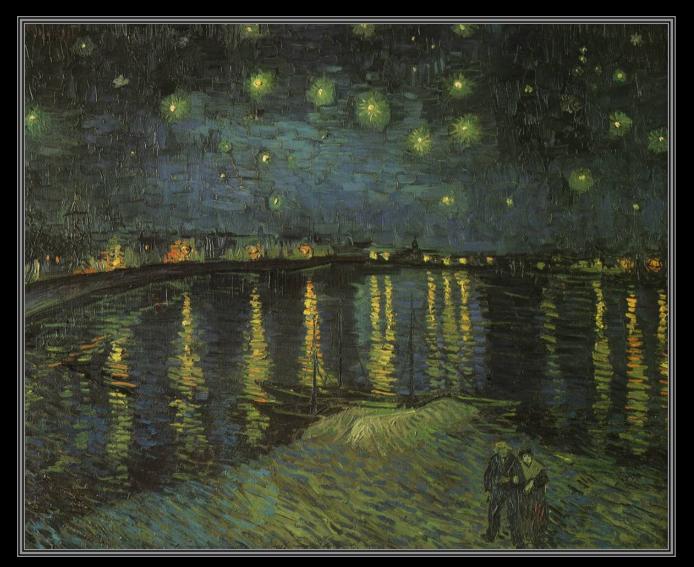
Copyright © 2004 Pearson Education, publishing as Addison Wesley.



The Local Supercluster



The Sky Is ... Dark?

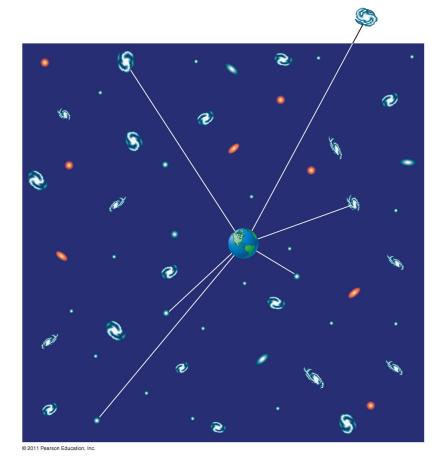


<u>Olber's Paradox</u>

Why is the night sky dark?

Imagine a universe:

- 1. Universe is finite in time
- 2. Universe is finite in space
- Universe is changing with time (and/or, either, or neither of the above)



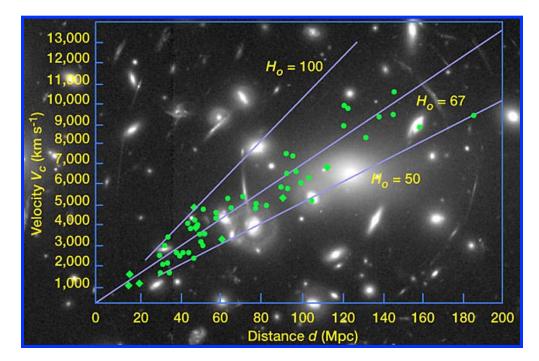
Beginnings in Cosmology

 Hubble's Law can be used to estimate the age of the universe

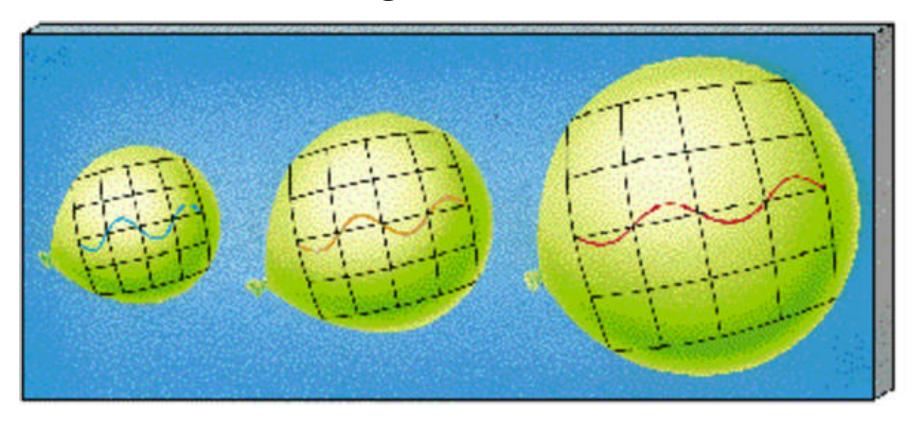
$$v_r = \frac{d}{t} = Hd$$

 $t = \frac{1}{H} \approx 1.5 \times 10^{10}$ years

 The universe had a beginning in time, and its age is around 15 billion years.

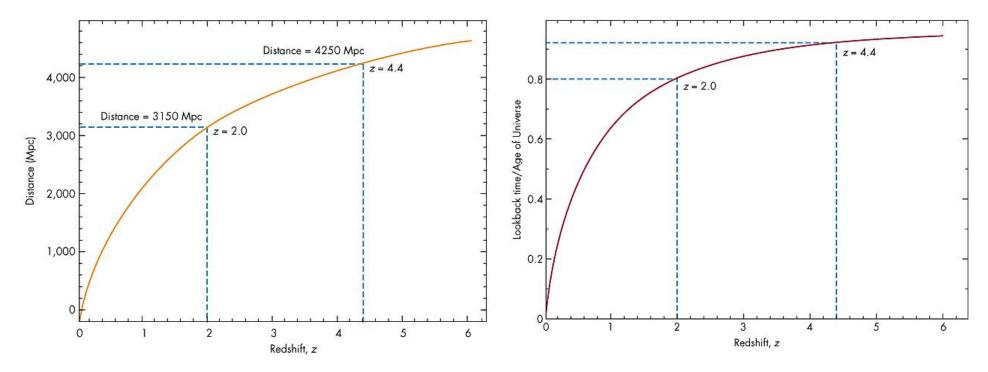


<u>Expansion and a New Look at</u> <u>Cosmological Redshift</u>



Observed redshifts of galaxies is not due to their motion but instead due to the literal expansion of space itself.

<u>Redshift as a Ruler</u>



The linear Hubble law only works for relatively nearby galaxies. At very high redshifts (z), the Hubble law becomes a curve. Above left: Distance with redshift. Above right: Look-back time with redshift.

The Cosmological Principle

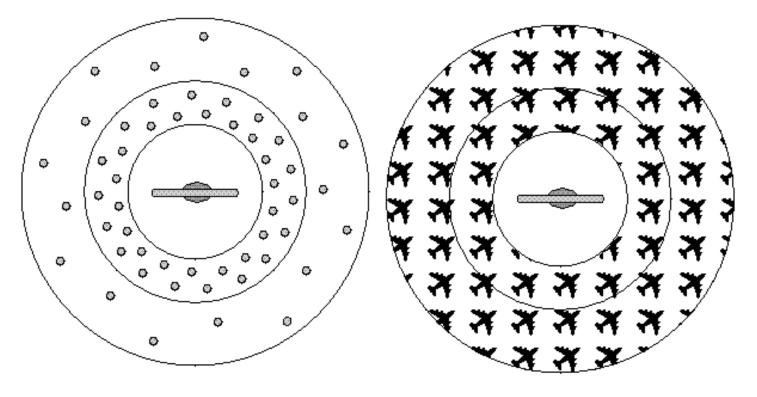
The universe is

- Homogeneous observers everywhere and anywhere see the "same" thing AND
- 2. Isotropic every direction presents the same view

These two assumptions form the basis for our interpretations of how the universe works.

[perfect cosmological principle includes "static"]

<u>Contrasting Homogeneity</u> <u>and Isotropy</u>



Is this *homogeneous* and *isotropic*? Which aspect is it not?

Outside the central sphere, is this universe *homogeneous* and *isotropic*? Which aspect is it not?

Special Relativity



- Einstein (1905): Two fundamental postulates
 - The laws of physics are the same for everyone and everywhere
 - The speed of light (300,000 km/s) is the same for everyone, even when things are in motion
- Relativity is a <u>misnomer</u>; the theory is one of <u>absolutism</u>, but an odd one in terms of the idea that the light speed is the same for all
- The counterintuitive aspect is that observers no longer agree on things being **simultaneous**

The Train Car Experiment

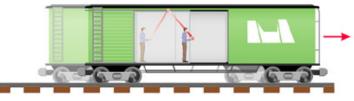
Inside the train, the ball goes up and down.



Outside the train, the ball appears to be going faster: It has the same up-and-down speed, plus the forward speed of the train.



The faster the train is moving, the faster the ball appears to be going to the outside observer.



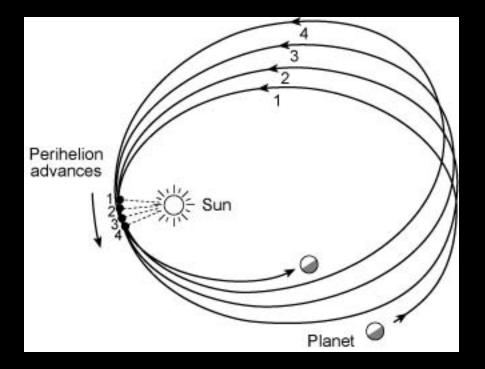
Copyright @ 2004 Pearson Education, publishing as Addison Wesley.

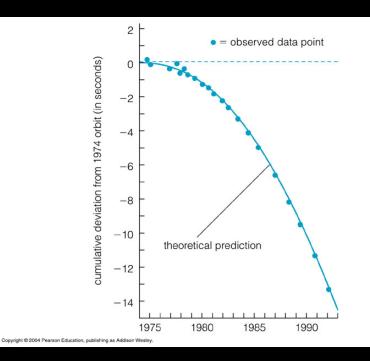
<u>General Relativity</u>

- Einstein (1915): A description of gravity relating matter and spacetime
- Spacetime 4D "space": location in time <u>and</u> regular 3D space (e.g., time, N/S, E/W, alt)
- Principle of Equivalence:

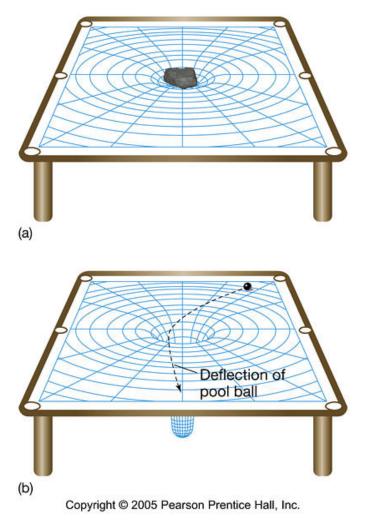
<u>Cannot</u> distinguish between a uniform gravitational field versus a frame of reference undergoing uniform acceleration.

Test of G.R.: Precession of Mercury

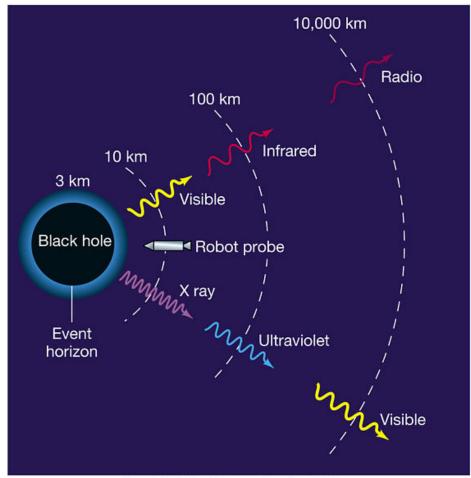




The Bizarre of G.R.



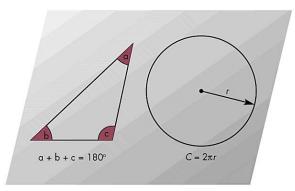
Mass distorts space

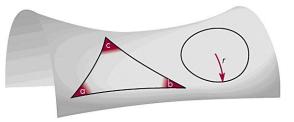


Copyright © 2005 Pearson Prentice Hall, Inc.

Gravitational redshift

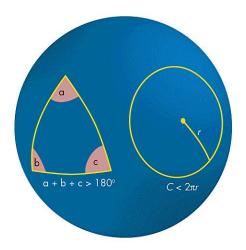
Curved Spacetime – Consequences



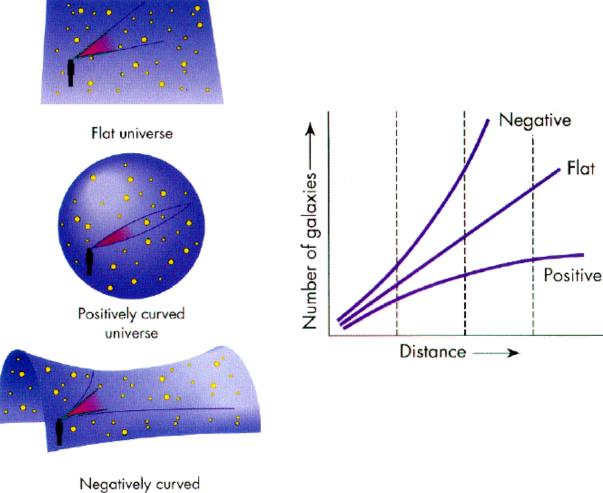


Triangle: $a + b + c < 180^{\circ}$ Circle: Circumference (C) > $2\pi r$

Examples for how the shape of a surface modifies familiar rules of geometry.



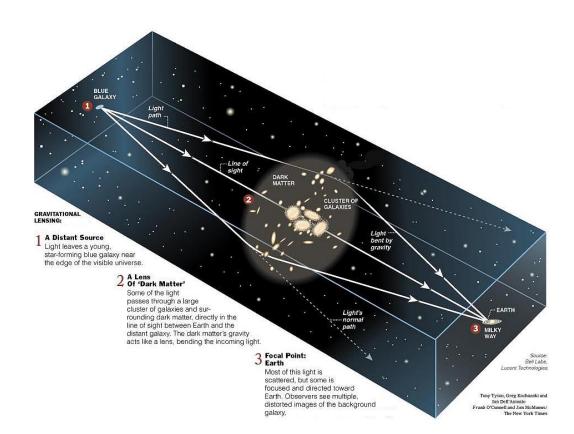
Measuring Curvature



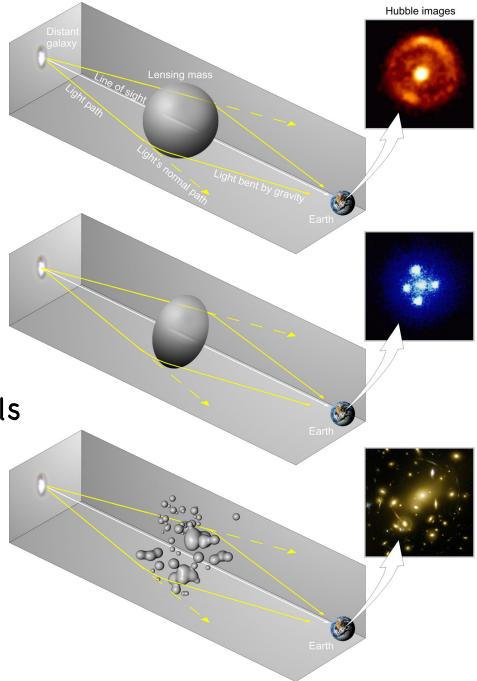
universe

Curvature of Spacetime

Einstein's theory predicts that matter "curves" spacetime. Consequently, light "bends" when passing near a mass. Phenomenon is "lensing" and is actually observed!

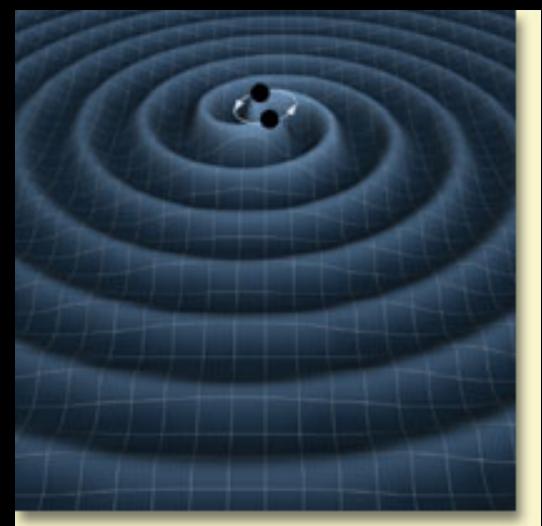


"Gravy" Lensing as a Tool



What you can see reveals what you cannot see: lensing can reveal mass distribution, especially dark matter

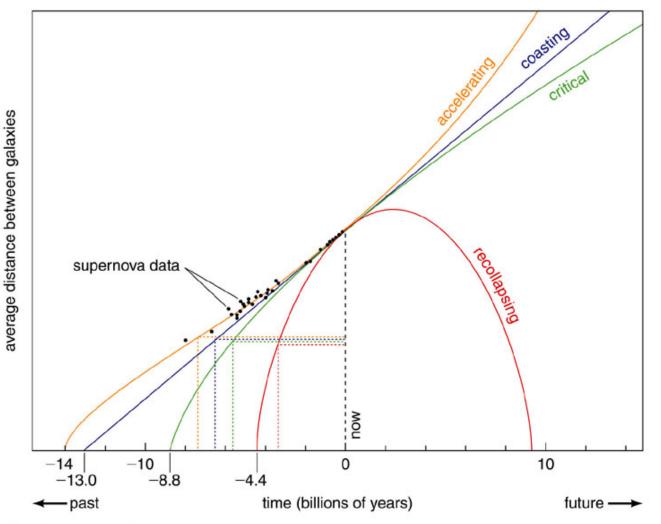
The Cutting Edge in G.R.: Gravitational Waves



Standard Big Bang Model

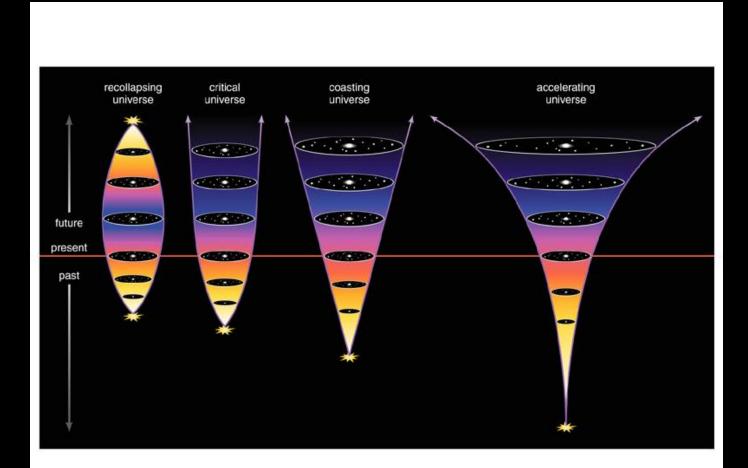
- Expansion suggests
 - a) A beginning
 - b) An "explosion"
- Big Bang highlights:
 - Universe was hot, dense, and expanding (at an age of 10⁻⁶ sec, temperature about 10¹³ K)
 - 2) A universe dominated by light. High energy photons are so energetic that they can combine to make basic matter: protons, neutrons, electrons.
 - 3) At t=100 sec, T ~ 10^{11} K, and nuclear fusion leads to the formation of He (and some other stuff)

Cosmological Models



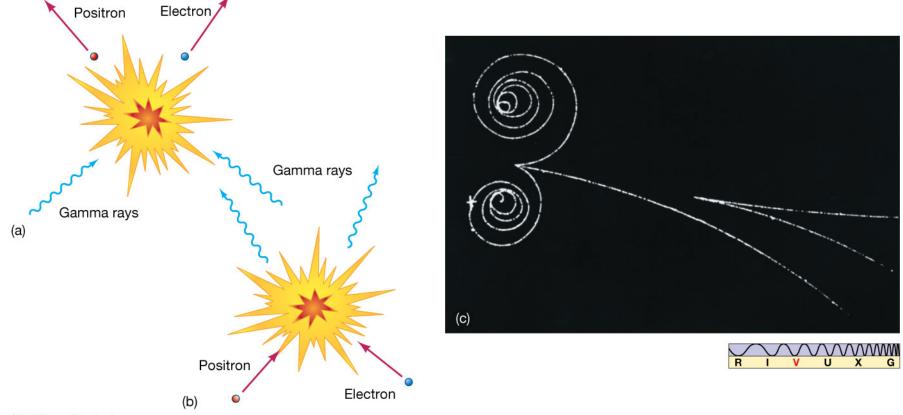
Copyright @ 2004 Pearson Education, publishing as Addison Wesley.

Expansion/Contraction of the Universe with Time for Different Models



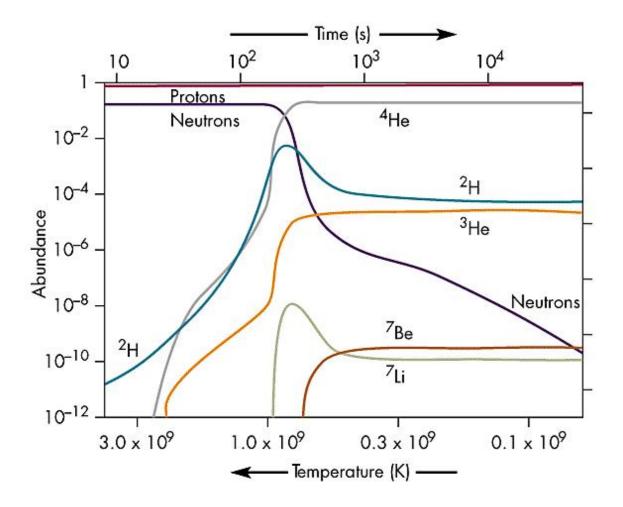
Copyright © 2004 Pearson Education, publishing as Addison Wesley.

The High Energy Early Universe: Pair Production and Annihilation

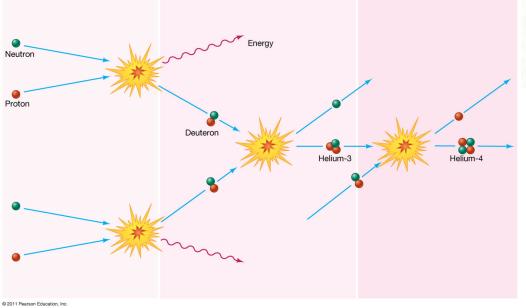


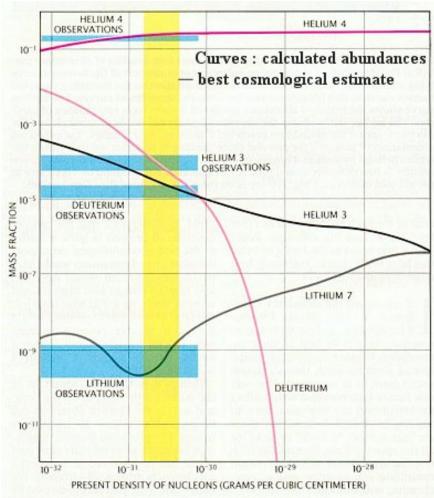
© 2011 Pearson Education, Inc.

Nucleosynthesis in the Early Universe: Predictions of the Big Bang Model



<u>Big Bang</u> Nucleosynthesis <u>Constraints</u>

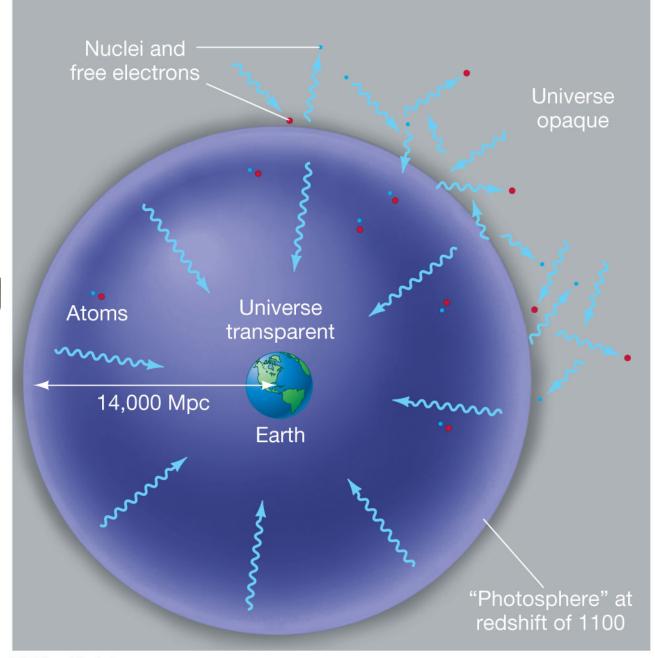




Big Bang Highlights (cont.)

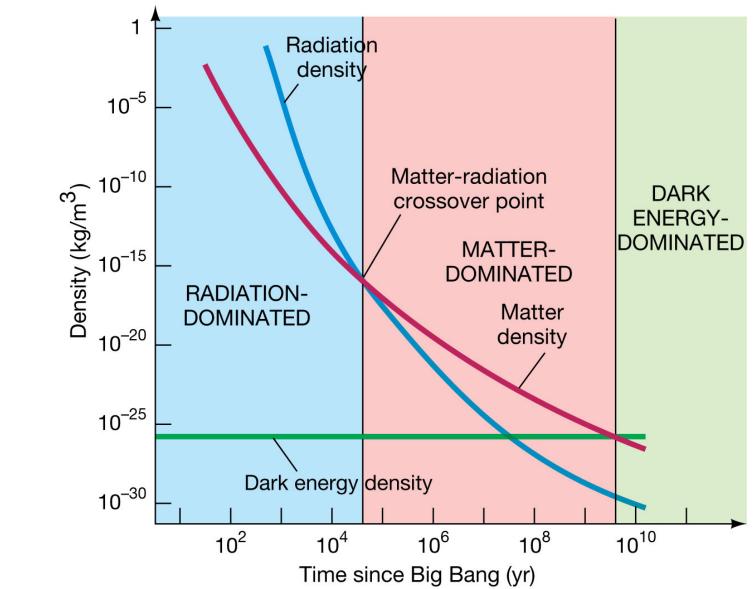
- At t ~ 700,000 yrs, T ~ 3000K, photons too weak to ionize H-atoms. Matter and light "decouple" (no longer interact)
- 5) At t ~ 1 billion yrs, galaxies begin to form (?)

Decoupling Era



© 2011 Pearson Education, Inc.

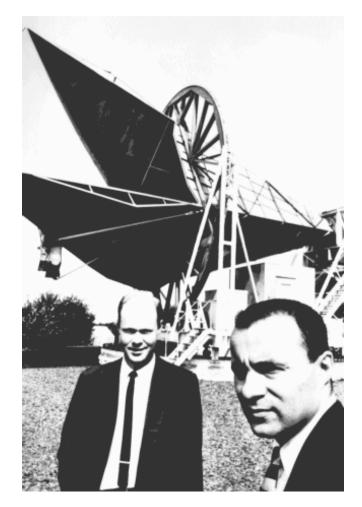
Universe Eras



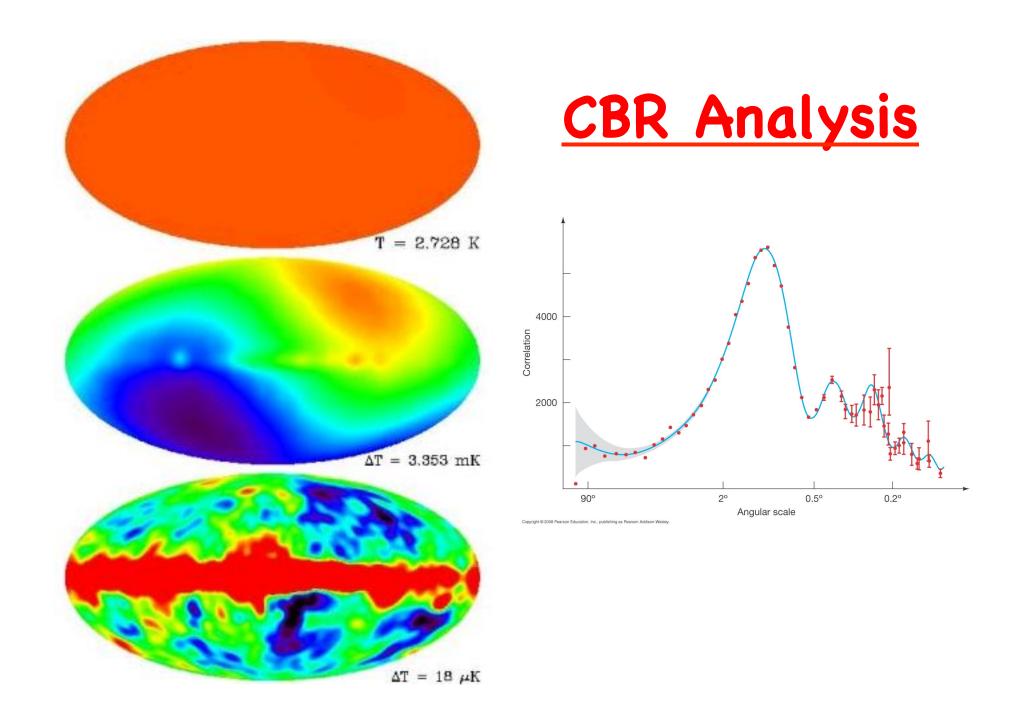
© 2011 Pearson Education, Inc.

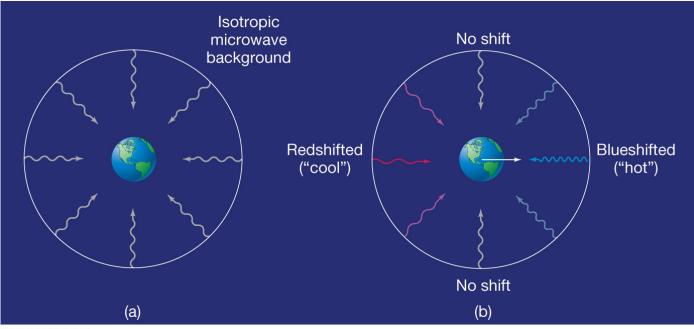
<u>Prediction and Discovery of the</u> <u>Cosmic Background Radiation (CBR)</u>

- 1940's, Gamow: predicts
 - a remnant glow from the early, hot universe should "pervade" space and be highly redshifted
 - BB glow of T \sim 5-50K
- 1964, Penzias and Wilson: stumbled upon CBR while looking for sources of radio interference



Found T=2.7K !

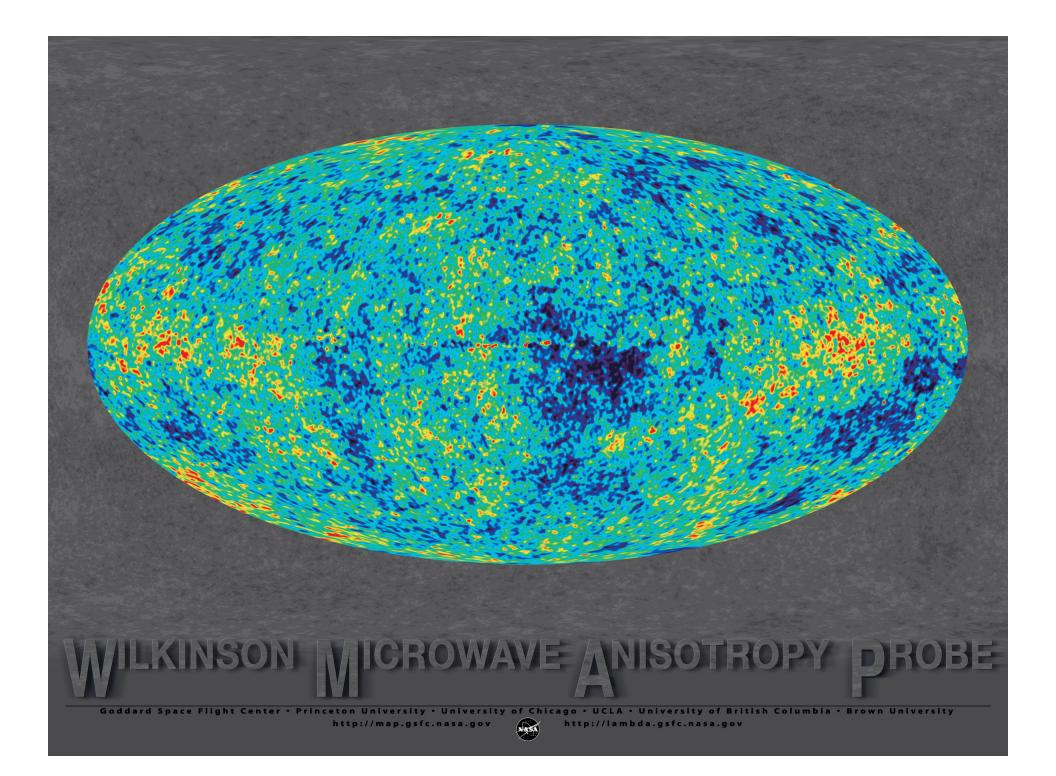


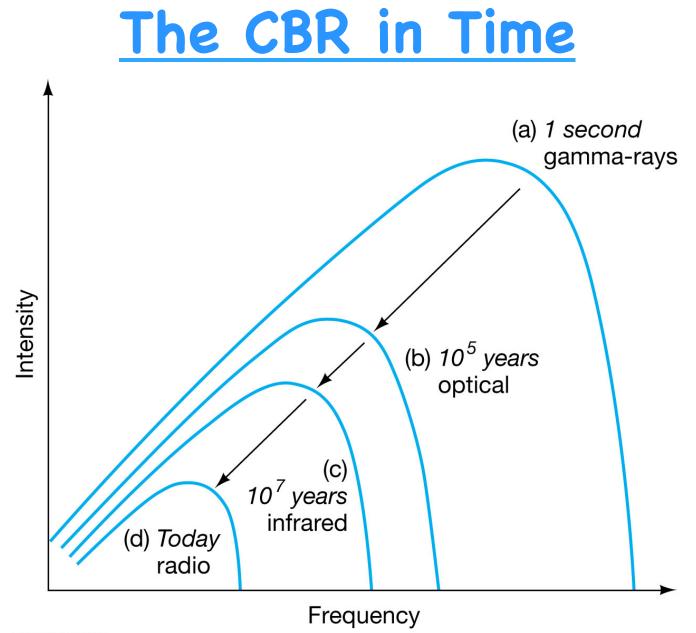


© 2011 Pearson Education, Inc.

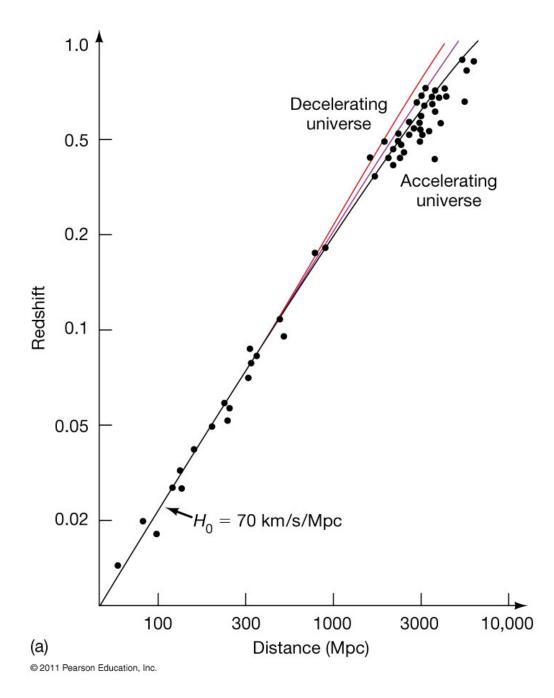
CBR: Dipole

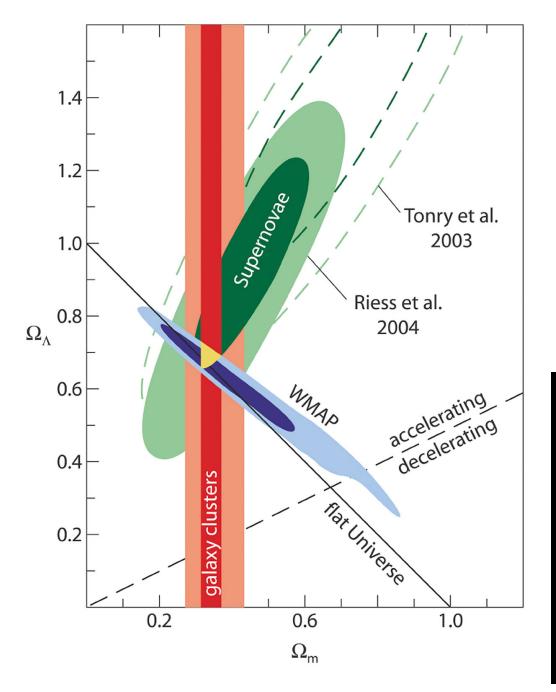
-4 mK +4 mK



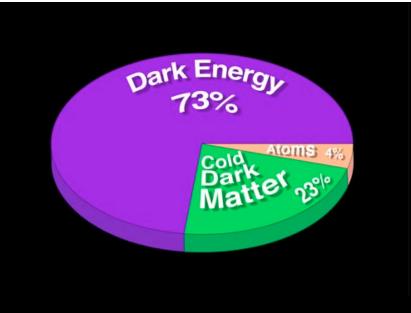


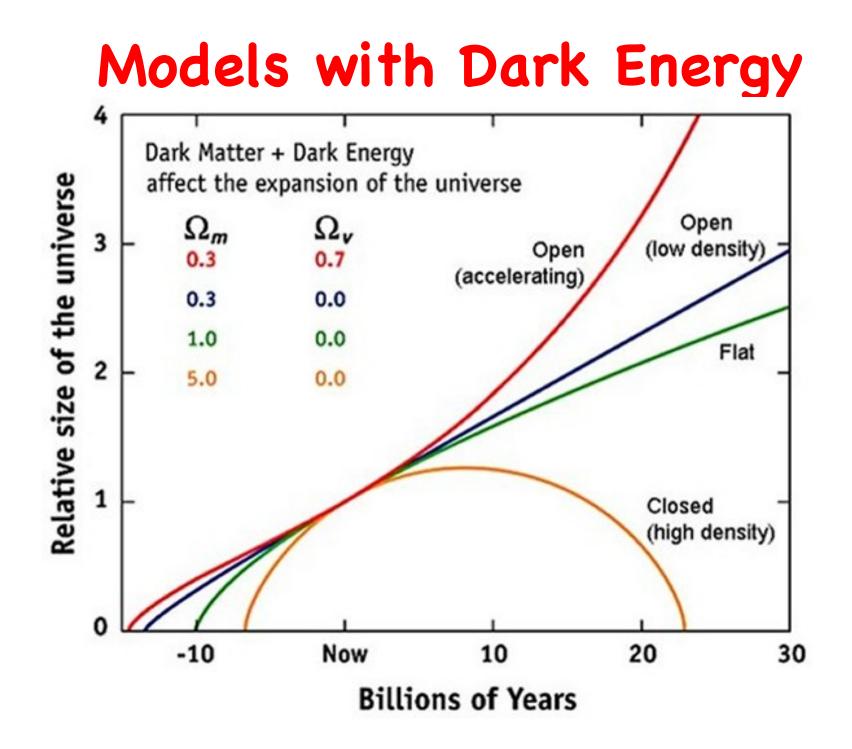
Evidence from Sne for an Accelerating Expansion of the Universe



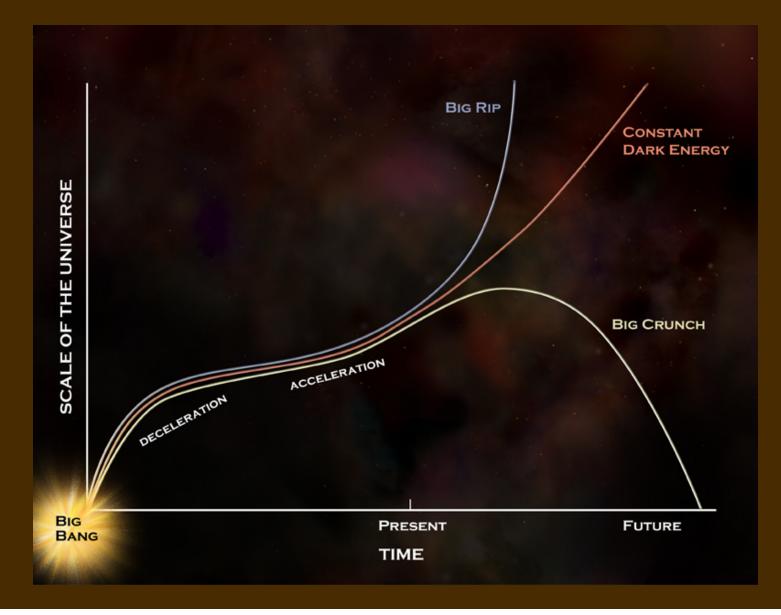


Narrowing the Scope of Cosmological Models





Cosmological Models with D.E.



<u>Oddities of the Universe</u>

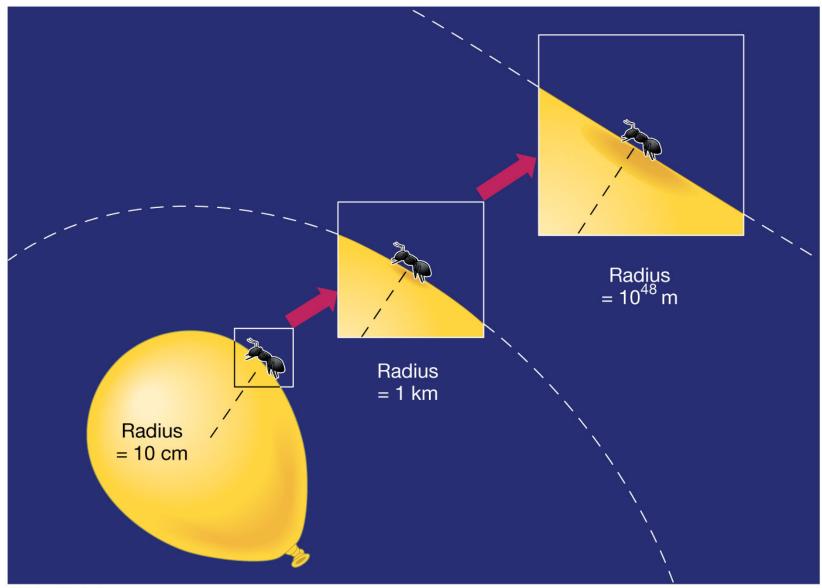
The universe shows oddities that the Big Bang model would not predict:

- Flatness problem the curvature of space could be positive or negative to any degree, or just exactly zero ("flat"); it appears to be very near zero; why that special value?!
- Horizon problem the cosmic background temperature is extremely uniform, but how can one side of the universe "know" about the temperature at the other side?!

Inflationary Theory

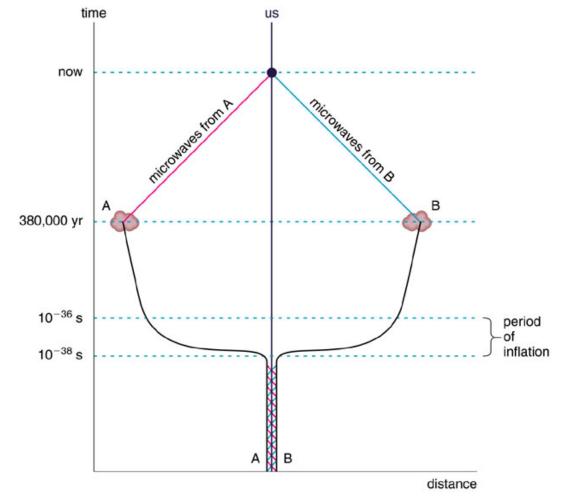
- This is a fairly recent theory that attempts to explain the oddities. It suggests an early, short period of **exponential** expansion of space.
- How does this help?
 - If you take a ballon and manage to expand it to the size of the Earth, then its surface will appear to be flat like the ground, even though you know that the surface is curved and closed. Rapid expansion makes any kind of initial curvature appear flat.
 - Also, rapid expansion suggests that regions vastly separated today were far more intimately connected in the past.
 - YES, we are talking about faster-than-light-speed expansion of space.
- Physical Motivation: Production of the major forces of nature. As if the universe experienced a paradigm shift and went ballistic. :)
- Evidence: It explains some oddities, plus details about the CBR fluctuations are consistent with predictions of inflationary theory.

Inflation and the Flatness Problem



© 2011 Pearson Education, Inc.

Graphic for Inflation



Copyright © 2004 Pearson Education, publishing as Addison Wesley.