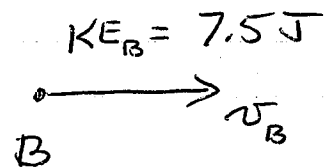
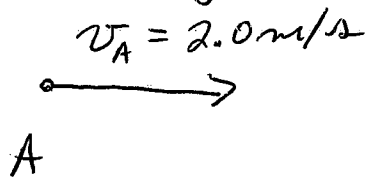


PHYS-2010  
GENERAL PHYSICS I  
DR. LUTERMOSER'S CLASS  
SOLUTION SET # 2  
TO THE SUPPLEMENTAL PROBLEMS

2-1

1.

$$m = 0.60 \text{ kg}$$



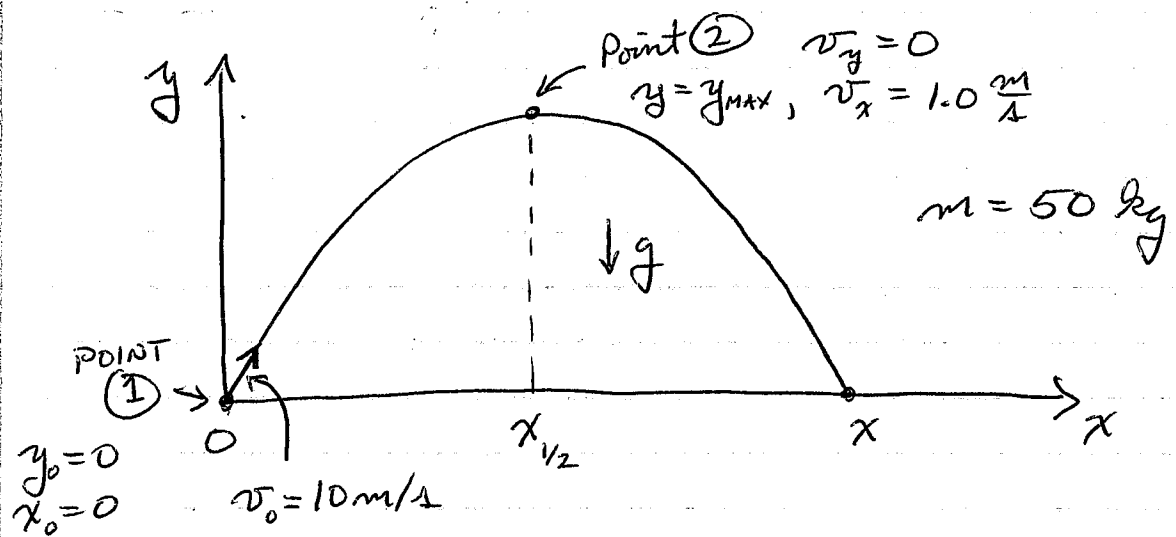
$$\begin{aligned} \text{a) } KE_A &= \frac{1}{2} m v_A^2 = \frac{1}{2} (0.60 \text{ kg}) (2.0 \frac{\text{m}}{\text{s}})^2 \\ &= \boxed{1.2 \text{ J}} \end{aligned}$$

$$\begin{aligned} \text{b) } KE_B &= \frac{1}{2} m v_B^2, & v_B^2 &= \frac{2(KE_B)}{m} \\ v_B^2 &= \frac{2(7.5 \text{ J})}{0.60 \text{ kg}} = \frac{15. \text{ kg m}^2/\text{s}^2}{0.60 \text{ kg}} \\ &= 25 \text{ m}^2/\text{s}^2 \end{aligned}$$

$$\boxed{v_B = 5.0 \text{ m/s}}$$

$$\text{c) } W = KE_B - KE_A = 7.5 \text{ J} - 1.2 \text{ J} = \boxed{6.3 \text{ J}}$$

2.



We could solve this with 2-D equations of motion, but here, it is easier to use conservation of energy:

$$KE_1 + PE_1 = KE_2 + PE_2$$

$$\frac{1}{2} m v_0^2 + m g y_0 = \frac{1}{2} m v^2 + m g y_{MAX}$$

At  $y_{MAX}$  :  $v = \sqrt{v_x^2 + v_y^2} = \sqrt{(1.0 \frac{m}{s})^2 + 0^2}$   
 $= 1.0 \text{ m/s}$

and

$m g y_0 = 0$  since  $y_0 = 0$  (the ground)

so

$$\frac{1}{2} v_0^2 + g y_0 = \frac{1}{2} v^2 + g y_{MAX}$$

$$g y_{MAX} = \frac{1}{2} (v_0^2 - v^2)$$

$$y_{MAX} = \frac{v_0^2 - v^2}{2g} = \frac{(10 \text{ m/s})^2 - (1.0 \text{ m/s})^2}{2(9.80 \text{ m/s}^2)} = \boxed{5.1 \text{ m}}$$

3.

Running person dissipates  $0.60 \text{ J/kg/step} = W_R$

$m = 60 \text{ kg}$ ,  $P = 70 \text{ W}$ ,  $l_{\text{STEP}} = 1.5 \text{ m}$ ,  $v = ?$

$$\text{Power} = \frac{\text{work done}}{\Delta t} = \frac{W}{\Delta t}$$

$$W = (\text{work per step per unit mass}) (\text{mass}) (\# \text{ of steps}) \\ = W_R \times m \times N, \text{ where } N = \# \text{ of steps}$$

$$\text{As such, } P = \frac{W}{\Delta t} = \frac{W_R m N}{\Delta t} = W_R m \frac{N}{\Delta t}$$

$$\text{or } \frac{N}{\Delta t} = \frac{P}{W_R m} = \frac{70 \text{ W}}{(0.60 \frac{\text{J/step}}{\text{kg}})(60 \text{ kg})} = 1.9 \text{ steps/s}$$

Each step covers  $l_{\text{STEP}} = 1.5 \text{ m}$  or the total distance traveled is

$$\Delta x = N l_{\text{STEP}}$$

The running speed is then

$$\bar{v} = \frac{\Delta x}{\Delta t} = \frac{N l_{\text{STEP}}}{\Delta t} = \frac{N}{\Delta t} l_{\text{STEP}}$$

$$= (1.9 \text{ steps/s})(1.5 \text{ m/step})$$

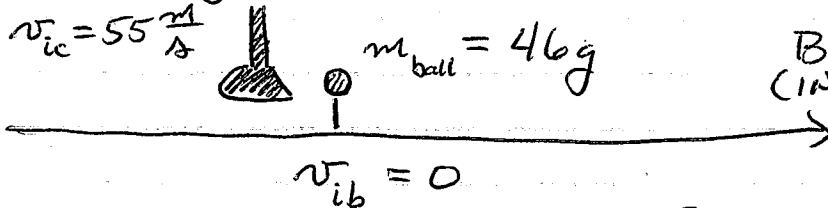
$$= \boxed{2.9 \text{ m/s}}$$

4.

$$m_{\text{club}} = 200 \text{ g}$$

$$v_{ic} = 55 \frac{\text{m}}{\text{s}}$$

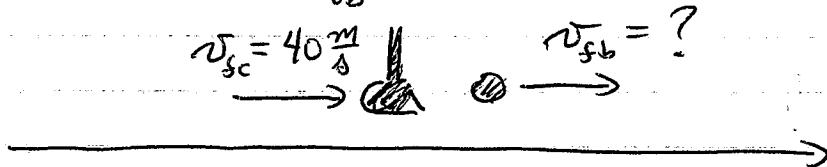
$$m_{\text{ball}} = 46 \text{ g}$$

BEFORE  
(INITIAL)

$$v_{ib} = 0$$

$$v_{sc} = 40 \frac{\text{m}}{\text{s}}$$

$$v_{fb} = ?$$

AFTER  
(FINAL)

Conservation of momentum:

$$m_c v_{ic} + m_b v_{ib} = m_c v_{fc} + m_b v_{fb}$$

$$m_b v_{fb} = m_c v_{ic} - m_c v_{fc}$$

$$v_{fb} = \frac{m_c}{m_b} (v_{ic} - v_{fc})$$

$$= \frac{200 \text{ g}}{46 \text{ g}} \left( 55 \frac{\text{m}}{\text{s}} - 40 \frac{\text{m}}{\text{s}} \right)$$

$$= \boxed{65 \text{ m/s}}$$

5.

The ratio of the kinetic energy of the Earth to that of the ball is

$$\frac{KE_{\oplus}}{KE_b} = \frac{\frac{1}{2} M_{\oplus} v_{\oplus}^2}{\frac{1}{2} m_b v_b^2} = \left( \frac{M_{\oplus}}{m_b} \right) \left( \frac{v_{\oplus}}{v_b} \right)^2$$

From conservation of momentum,

$$p_f = p_i = 0, \text{ giving } M_{\oplus} v_{\oplus} + m_b v_b = 0, \text{ or}$$

$$\boxed{\frac{v_{\oplus}}{v_b} = -\frac{m_b}{M_{\oplus}}}$$

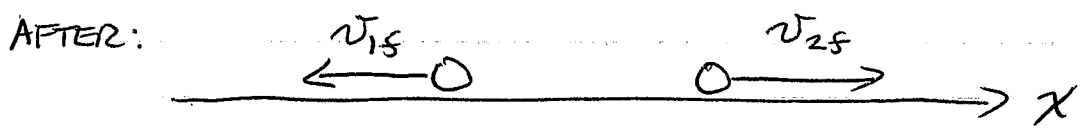
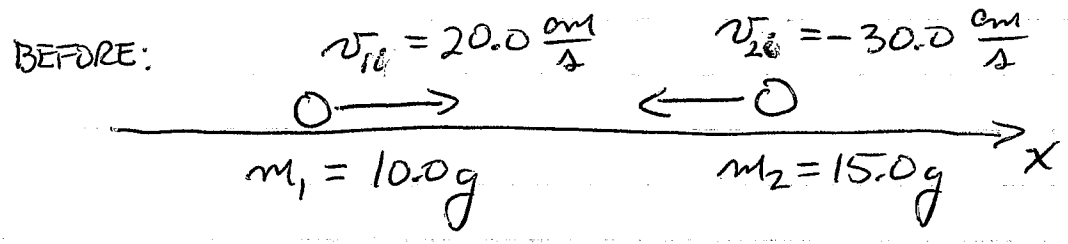
and then

$$\frac{KE_{\oplus}}{KE_b} = \left( \frac{M_{\oplus}}{m_b} \right) \left( -\frac{m_b}{M_{\oplus}} \right)^2 = \boxed{\frac{m_b}{M_{\oplus}}}$$

Using order of magnitude numbers:

$$\frac{KE_{\oplus}}{KE_b} = \frac{m_b}{M_{\oplus}} \approx \frac{1 \text{ kg}}{10^{25} \text{ kg}} \approx \boxed{10^{-25}}$$

6.



(1) Conservation of momentum:

$$m_1 v_{1i} + m_2 v_{2i} = m_1 v_{1f} + m_2 v_{2f}$$

(2) Conservation of kinetic energy (2nd form):

$$v_{1i} - v_{2i} = -(v_{1f} - v_{2f})$$

Now, using (1) gives

$$10.0 \text{g} (20.0 \text{cm/s}) + 15.0 \text{g} (-30.0 \text{cm/s}) = 10.0 \text{g} (v_{1f}) + 15.0 \text{g} (v_{2f})$$

$$200. \text{g cm/s} - 450 \text{g cm/s} = (10.0 \text{g}) v_{1f} + (15.0 \text{g}) v_{2f}$$

$$-250. \text{g cm/s} = (10.0 \text{g}) v_{1f} + (15.0 \text{g}) v_{2f}$$

(1a)  $-25.0 \frac{\text{cm}}{\text{s}} = v_{1f} + 1.50 v_{2f}$

and using (2) gives

$$20.0 \frac{\text{cm}}{\text{s}} - (-30.0 \frac{\text{cm}}{\text{s}}) = -(v_{1s} - v_{2s})$$

$$(2a) \quad 50.0 \text{ cm/s} = -v_{1s} + v_{2s}$$

Add (1a) & (2a):

$$\begin{array}{r} -25.0 \text{ cm/s} = v_{1s} + 1.50 v_{2s} \\ + 50.0 \text{ cm/s} = -v_{1s} + v_{2s} \\ \hline \end{array}$$

$$25.0 \text{ cm/s} = 0 + 2.50 v_{2s}$$

$$v_{2s} = \frac{25.0 \text{ cm/s}}{2.50} = \boxed{10.0 \text{ cm/s}}$$

Plugging this value into (2a):

$$50.0 \text{ cm/s} = -v_{1s} + 10.0 \text{ cm/s}$$

$$v_{1s} = 10.0 \text{ cm/s} - 50.0 \text{ cm/s}$$

$$= \boxed{-40.0 \text{ cm/s}}$$

7.

$\omega_0 = 0$ , after  $t = 3.20 \text{ s}$ ,  $\omega = 2.51 \times 10^4 \frac{\text{rev}}{\text{min}}$   
 $t_0 = 0$  (under constant  $\alpha$ )

a) Find angular acceleration:

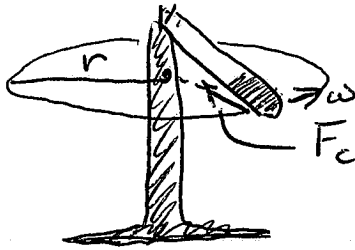
$$\begin{aligned} \alpha = \bar{\alpha} &= \frac{\omega - \omega_0}{t - t_0} = \frac{2.51 \times 10^4 \text{ rev/min} - 0}{3.20 \text{ s} - 0} \\ &= 7.84 \times 10^3 \frac{\text{rev}}{\text{min} \cdot \text{s}} \times \frac{1 \text{ min}}{60 \text{ s}} = 131 \text{ rev/s}^2 \\ &= 131 \frac{\text{rev}}{\text{s}^2} \times \frac{2\pi \text{ rad}}{1 \text{ rev}} = \boxed{821 \text{ rad/s}^2} \end{aligned}$$

b) Total angle turned during this time?

$$\begin{aligned} \theta &= \omega_0 t + \frac{1}{2} \alpha t^2 = 0 + \frac{1}{2} (821 \frac{\text{rad}}{\text{s}^2}) (3.20 \text{ s})^2 \\ &= \boxed{4.21 \times 10^3 \text{ rad}} = 2.41 \times 10^5 \text{ degrees} \end{aligned}$$



8.



$$m_{\text{cell}} = 3.0 \times 10^{-16} \text{ kg}$$

$F_c = 4.0 \times 10^{-11} \text{ N}$  toward center of circle

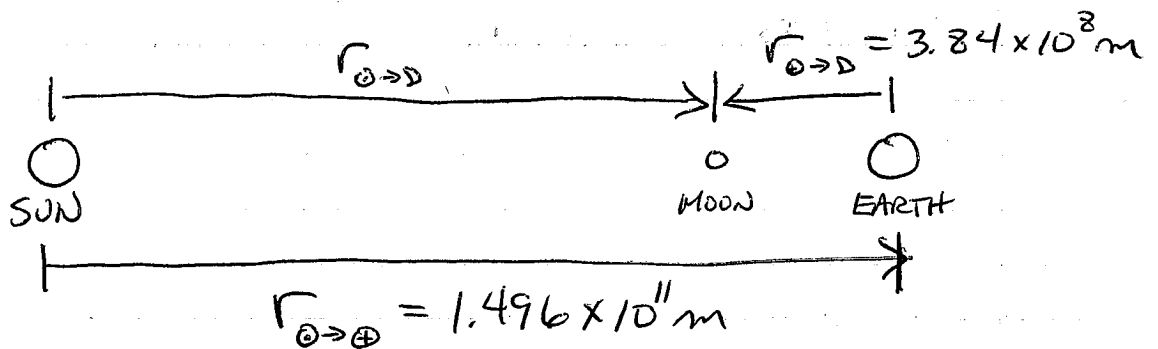
$$r = 0.150 \text{ m} \quad \left( = \frac{15.0 \text{ cm}}{100 \text{ cm/m}} \right)$$

$$F_c = m \frac{v_t^2}{r} = m r \omega^2$$

$$\omega = \sqrt{\frac{F_c}{m r}} = \sqrt{\frac{4.0 \times 10^{-11} \text{ N}}{(3.0 \times 10^{-16} \text{ kg})(0.150 \text{ m})}}$$

$$= 9.4 \times 10^2 \frac{\text{rad}}{\text{s}} \times \frac{1 \text{ rev}}{2\pi \text{ rad}} = \boxed{1.5 \times 10^2 \frac{\text{rev}}{\text{s}}}$$

9.



$$r_{\text{sun} \rightarrow \text{moon}} = r_{\text{sun} \rightarrow \text{earth}} - r_{\text{moon} \rightarrow \text{earth}} = 1.496 \times 10^{11} \text{ m} - 3.84 \times 10^8 \text{ m} = 1.492 \times 10^{11} \text{ m}$$

$$a) F_{\text{sun} \rightarrow \text{moon}} = \frac{G M_{\text{sun}} M_{\text{moon}}}{r_{\text{sun} \rightarrow \text{moon}}^2} = \frac{(6.673 \times 10^{-11} \text{ N m}^2/\text{kg}^2)(7.36 \times 10^{24} \text{ kg}) \times (1.991 \times 10^{30} \text{ kg})}{(1.492 \times 10^{11} \text{ m})^2}$$

$$\boxed{F_{\text{sun} \rightarrow \text{moon}} = 4.39 \times 10^{20} \text{ N}}$$

(2-10)

$$b) F_{\oplus \rightarrow D} = \frac{GM_{\oplus} M_D}{r_{\oplus \rightarrow D}^2} = \frac{(6.673 \times 10^{-11} \text{ Nm}^2/\text{kg}^2) \times (5.98 \times 10^{24} \text{ kg})(7.36 \times 10^{22} \text{ kg})}{(3.84 \times 10^8 \text{ m})^2}$$

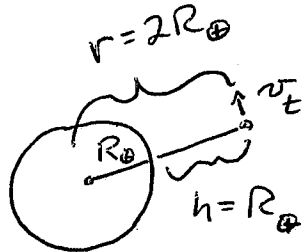
$$F_{\oplus \rightarrow D} = 1.99 \times 10^{20} \text{ N}$$

$$c) F_{\oplus \rightarrow \oplus} = \frac{GM_{\oplus} M_{\oplus}}{r_{\oplus \rightarrow \oplus}^2} = \frac{(6.673 \times 10^{-11} \text{ Nm}^2/\text{kg}^2) (5.98 \times 10^{24} \text{ kg})(5.98 \times 10^{24} \text{ kg})}{(1.496 \times 10^{11} \text{ m})^2}$$

$$F_{\oplus \rightarrow \oplus} = 3.55 \times 10^{22} \text{ N}$$

10.

a)



Set gravitational force = centripetal force

$$m = \text{satellite mass} = 600 \text{ kg}$$

$$F_c = F_g$$

$$m \frac{v_t^2}{r} = \frac{GM_{\oplus} m}{r^2}$$

$$v_t = \sqrt{\frac{GM_{\oplus}}{r}} = \sqrt{\frac{GM_{\oplus}}{2R_{\oplus}}}$$

$$= \sqrt{\frac{(6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2) (5.98 \times 10^{24} \text{ kg})}{2(6.38 \times 10^6 \text{ m})}}$$

$$= 5.59 \times 10^3 \text{ m/s}$$

b) The period of the satellite's motion is

$$T = \frac{2\pi r}{v_t} = \frac{2\pi [2(6.38 \times 10^6 \text{ m})]}{5.59 \times 10^3 \text{ m/s}}$$

$$= \boxed{1.43 \times 10^4 \text{ s}} \times \frac{1 \text{ hr}}{3600 \text{ s}} = \boxed{3.98 \text{ hr}}$$

c) The gravitational force acting on the satellite:

$$F = \frac{GmM_{\oplus}}{r^2} = \frac{GmM_{\oplus}}{4R_{\oplus}^2}$$

$$= \frac{(6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2)(600 \text{ kg})(5.98 \times 10^{24} \text{ kg})}{4(6.38 \times 10^6 \text{ m})^2}$$

$$= \boxed{1.47 \times 10^3 \text{ N}}$$