Chapter 9. The Moon



Figure 9.4. An Apollo mission photo of the Moon showing part of the near side (on the left) and part of the far side (on the right).

Note. In this section we survey physical properties of the Earth's moon. The Moon has has a *synchronous rotation* and the same side of the moon always faces the Earth.

Note. The albedo of the Moon is 7%. That is, 7% of the light that strikes the surface of the moon is reflected back.

Note. There is no atmosphere on the Moon, so features such as *craters* do not erode. Other features include *rays* (which emanate from craters) and *rilles* (valleys like canyons). The Moon's mass is 1.2% of the Earth's mass and has 16.5% of Earth's surface gravity.

Note. The Moon's surface is dominated by *highlands* (mountainous terrain), *maria* (large, smooth areas that are darker than the surrounding areas). The near side of the Moon is dominated by maria, while the far side is "rough" overall.

Note. Apollo missions have returned 800 pounds of lunar samples. The linear soil (or *regolith*) is loosely packed rock fragments and glassy minerals, probably created by meteor impacts. There are *breccias* which are rock fragments cemented together, and *basalts* which are igneous silicates common from lava flows. The relative abundance of several atomic isotopes is the same for the Earth and Moon, although this varies throughout the Solar System.

Note. Moonquakes were detected by Apollo mission equipment. The regolith is 10m thick and on a thicker layer of loose rubble. The crust is 50 km to 100 km thick, maybe thicker on the far side. Next is the mantle with a rigid lithosphere and a semi-liquid astherosphere. The inner 500 km is a dense core. The Moon is 1740 km in radius. See Figure 9.17.

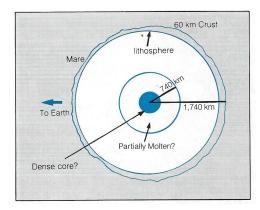


Figure 9.17. The internal structure of the Moon.

No magnetic field was detected, indicating a core that is probably not molten. Also, no evidence of present tectonics was detected.

Note. The hypotheses on the origin of the Moon include:

- (1) The Moon and Earth formed together (*coeval formation* or *binary accretion*). This is supported by the compositions of the Earth and Moon, but is mechanically difficult to explain.
- (2) The Moon is material split off from the Earth during formation. This is also supported by compositions, by why did the Moon split off? A fast rotation rate would be necessary of the Moon to just spin off.
- (3) The Moon formed elsewhere and was captured. But this doesn't explain the similar composition of the Earth and Moon, and this presents mechanical problems also.
- (4) The Moon formed when a huge Mars-sized planetesimal hit Earth and knocked it off. This is supported by the compositions and there are known to have been many planetesimals in the early solar system.

Note. The Moon's small gravitational pull allows gases to escape and it has no atmosphere. The maria are results of huge impacts. The Earth and Moon were once closer; tidal forces have cause slowing of the Moon's orbital speed and separation of the Earth and Moon.

Name	Nation	Year	Results
Luna 2	USSR	1959	impacted surface
Luna 3	USSR	1959	photographed the far side
Ranger 7	USA	1964	photographed surface, impacted
Luna 9	USSR	1966	sent photos from surface
Luna 10	USSR	1966	orbiter
Surveyor 1	USA	1966	soft landing
Lunar Orbiter 1	USA	1966	orbiter, photographed surface
Zond 5	USSR	1968	round trip with life forms aboard
Apollo 8	USA	1968	round trip with humans
Apollo 11	USA	1969	first humans land on Moon
Luna 16	USSR	1970	returned a sample form surface

Note. Satellites launched to the Moon include:

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