#### STUDY GUIDE FOR TEST IV

# CHAPTER 25: STRUCTURE AND ORGANIZATION OF THE GALAXY

- 1. nucleus
- 2. halo
- 3. dimensions of the Galaxy
- 4. nature and use of Cepheid variables
- 5. RR Lyrae variables
- 6. Kapteyn's result and the reason for it
- 7. Shapely's result and method
- 8. the Sun's position in the Galaxy
- 9. rotation of the Galaxy
- 10. peculiar velocity
- 11. distribution of O and B stars
- 12. hydrogen clouds and the 21 cm emissions
- 13. mass determination from the Sun's orbit
- 14. number of stars
- 15. types of stars near center of the Galaxy
- 16. implications of high speed clouds near nucleus
- 17. evidence for a black hole at the center of the Galaxy
- 18. age of globular clusters
- 19. amount of matter thought possibly to be in the halo

#### CHAPTER 26: THE INTERSTELLAR MEDIUM

- 1. two types of interstellar material
- 2. two effects of dust grains
- 3. interstellar extinction
- 4. scattering and its effect on color
- 5. use of B-V color index to determine amount of dust in the line of sight of a star
- 6. alignment of grains and the Galaxy's magnetic field
- 7. origins of interstellar grains
- 8. detection of interstellar gas
- 9. diffuse clouds (size and mass)
- 10. emission nebula
- 11. H II regions
- 12. reflection nebula and its color
- 13. dark clouds and the detection of molecules
- 14. organic molecules in dark clouds

15. movements of diffuse clouds and effect of supernova and stellar winds

16. superbubbles

## CHAPTER 27: THE FORMATION AND EVOLUTION OF THE GALAXY

- 1. Population I stars, Population II stars, and the distribution of heavy elements
- 2. disk population
- 3. abundance gradient
- 4. high velocity stars
- 5. "cycling" of interstellar material
- 6. density wave theory
- 7. effect of density wave on distribution of young stars
- 8. possible cause of density wave
- 9. significance fo the distribution of Pop II stars and globular clusters to the early structure of the Galaxy
- 10. starburst galaxy

## CHAPTER 28. THE NATURE OF THE NEBULAE

- 1. elliptical galaxies and subclasses
- 2. spiral galaxies and subclasses
- 3. barred spirals
- 4. tuning fork diagram
- 5. Type I and Type II irregulars
- 6. Cepheid variables as distance estimators
- 7. brightest stars and supernovae as distance estimators
- 8. apparent diameter as distance estimator
- 9. brightest galaxies in a cluster as distance estimators
- 10. Tully-Fisher method
- 11. rotational velocity and rotation curve in mass determination
- 12. velocity dispersion and its cause in elliptical galaxies
- 13. use of binary galaxies in mass determina-
- 14. dwarf ellipticals
- 15. mass-to-light ratio for elliptical and spiral galaxies

- 16. evolutionary reason for the shape of elliptical galaxies EXPLOSIVE N
- 17. role of halo in bar formation

#### CHAPTER 29: CLUSTERS AND SU-PERCLUSTERS: THE DISTRIBUTION OF THE GALAXIES

- 1. cluster
- 2. Local Group
- 3. Andromeda Galaxy
- 4. size of the Local Group
- 5. Magellanic Clouds
- 6. migration of large galaxies in rich clusters
- 7. effect of gravitational interactions on spiral galaxies
- 8. two methods for deterimining mass of a cluster
- 9. temperature and composition of intercluster gas
- 10. superclusters and their structure
- 11. top-down theory
- 12. bottom-up theory
- 13. dark matter
- 14. cause of filament structure of superclusters
- 15. cosmic strings

## CHAPTER 30: UNIVERSAL EXPANSION AND THE COSMIC BACKGROUND

- 1. Hubble's discovery/Hubble's Law
- 2. expansion of the universe
- 3. age determination of universe from H
- 4. determination of distance of a galaxy from its redshift
- 5. big bang
- 6. Gamow's prediction
- 7. discovery of Penzias and Wilson and its significance
- 8. microwave background/3-degree background radiation and its cause
- 9. isotropy of microwave background
- 10. movement of Local Group relative to Virgo Cluster
- 11. streaming motion and the "great attractor"

## CHAPTER 31: PECULIAR GALAXIES, EXPLOSIVE NUCLEI, AND QUASARS

- 1. radio galaxies
- 2. radio lobes and ejected jets
- 3. ejected gas from galactic cores
- 4. production of jets
- 5. Seyfert galaxies
- 6. quasars and their redshifts
- 7. BL lac objects
- 8. time when quasars lived
- 9. radio and X-ray emissions by quasars
- 10. jets from quasars
- 11. light variation of quasars
- 12. absorption lines in quasar spectrum and two possible causes
- 13. probable relationship of quasars, Seyfert glaxies, and normal galaxies
- 14. possible power source for quasars

#### CHAPTER 32: COSMOLOGY: PAST PRESENT, AND FUTURE OF THE UNI-VERSE

- 1. cosmology
- 2. cosmogony
- 3. "underlying assumptions" (applications of physical laws, homogeneity, isotropy)
- 4. Cosmological Principle
- 5. general theory of relativity and its implications for moving objects, space, and the expanding universe
- 6. curvatures of the universe (open universe, closed universe, flat universe)
- 7. two approaches to testing for openness or closedness
- 8. methods for estimating total mass and results
- 9. places where dark matter may be
- 10. methods for estimating deceleration
- 11. inflationary universe and its advantages
- 12. history of the first few minutes of the universe
- 13. time of formation of galaxies and clusters
- 14. scenario for the two possible endings of the universe