History of Geometry

Chapter 5. Trigonometry

5.10. The Great Discoveries of Kepler and Newton-Proofs of Theorems



1 Theorem 5.8 (Newton's Theorem 1, Principia Mathematica)

2 Newton's Lemma

3 Theorem 5.9. (Proposition 11 of Newton's Principia)

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Proof. We go through the argument given in Ostermann and Wanner (which is the same as given by Newton in *Principia*). The argument is a discretized version of the physical problem. Newton appeals to another result (Lemma III in his Book I) to move from the discrete to the continuous case.

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Theorem 5.8 (Newton's Theorem 1, continued 1)

Proof (continued). Next, we let the force act again at point *B* by an amount $f \Delta t$. Now consider the triangles *ABS* and *BcS* in the figure. We claim that these triangles have the same areas. We argue this below.



The Figure for Newton's Theorem 1 from Motte's 1846 translation of *Principia*.

Theorem 5.8 (Newton's Theorem 1, continued 2)

Proof (continued). To see that triangles ABS and BcS have the same area, reflect triangle ABS about the line containing points S and B and let A' denote the image of A. We then see that the triangle have a common base SB and the same altitude. So the areas are the same (by Euclid I.41).



Theorem 5.8 (Newton's Theorem 1, continued 3)

Proof (continued). Next, we let the force affect the object at point *B*. By Newton's Second Law (or Lex 2), the force acts along the line from *B* to *S*; the "change of motion" (as it's called in Lex 2) is represented as the vector from *B* to *V* (we denote this vector, and similar vectors, as \overrightarrow{BV}) in Figure 5.29 (right) below. Adding this change in motion to the velocity of the object before the application of the force (represented by the vector \overrightarrow{AB} in the figure; think of it as the change in position over time Δt), we get the new velocity by adding vectors

 \overrightarrow{AB} and \overrightarrow{BV} to get the resultant vector \overrightarrow{AV} . We translate \overrightarrow{AV} to point *B* to get velocity vector \overrightarrow{BC} .



Theorem 5.8 (Newton's Theorem 1, continued 3)

Proof (continued). Next, we let the force affect the object at point *B*. By Newton's Second Law (or Lex 2), the force acts along the line from Bto S; the "change of motion" (as it's called in Lex 2) is represented as the vector from B to V (we denote this vector, and similar vectors, as \overrightarrow{BV}) in Figure 5.29 (right) below. Adding this change in motion to the velocity of the object before the application of the force (represented by the vector \overrightarrow{AB} in the figure; think of it as the change in position over time Δt), we get the new velocity by adding vectors AB and BV to get the resultant vector \overrightarrow{AV} . We translate \overrightarrow{AV} to point B to get velocity vector \overrightarrow{BC} .

Theorem 5.8 (Newton's Theorem 1, continued 4)

Proof (continued). By construction above, vector \overrightarrow{AB} equals vector \overrightarrow{BC} . So when translating vectors \overrightarrow{AB} and \overrightarrow{AV} to point *B*, we see that segment *Cc* is parallel to segment *SB*. So the area of triangles *BCS* and *BcS* are the same since they have the same base, *SB*, and the same heights (represented, respectively, in the figure by the segments with the right-angle symbols which end at *C* and *c*; we are again using Euclid I.41 here).



Theorem 5.8 (Newton's Theorem 1, continued 5)

Proof (continued). Since, as shown above, the triangles *ABS* and *BcS* have the same areas, then we have that triangles *ABS* and *BCS* have the same areas. Similar, the areas of triangles *ABS*, *BCS*, *CDS*, *DES*, etc. are all the same.

So the claim holds when we have the discrete version of the problem with the force applied as impulses at equal time steps Δt .



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In the spirit of Calculus 1 (MATH 1910), we could consider the area swept out by the line segment joining the Sun and the orbiting object as a function time A(t). We then have from Newton's discrete argument that $\Delta A/\Delta t$ is a constant. So we have $\lim_{\Delta t \to 0} \frac{\Delta A}{\Delta t} = \frac{dA}{dt}$ is constant (this is the constant of proportionality in Newton's theorem).

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Newton's Lemma

Newton's Lemma. Let APQ be an ellipse with focus S and suppose P to be the position of the planet moving towards Q, while the point R moves on the tangent with S, Q, P collinear. Let T be the orthogonal projection of Q onto PS (see Figure 5.30, right). Then if the distance PQ tends to zero, we have $RQ \approx (\text{Constant}) \cdot QT^2$, where the constant is independent of the position of P on the ellipse.



Fig. 5.30. Reproductions from Newton's autograph (1684), manuscript Cambridge Univ. Lib. Add. 3965^6 ; the force acting on a moving body (left); picture for Newton's lemma (right). Reproduced by kind permission of the Syndics of Cambridge University Library

Proof.



Fig. 5.31. Newton's proof of his lemma

From this figure, we have by Apollonius II.6 (see Section 3.2. The Ellipse, Figure 3.7(b)) that the tangent *PR* is parallel to the "diameter" *DCK* which is conjugate to diameter *GCP*. Let the lengths of these diameters be 2*d* and 2*c*, respectively, as labeled in the Figure 5.31. Through focus *H* we draw a parallel to *DK* (light dotted line) to determine point of intersection *I* with segment *SP*. Through point *Q* we draw a parallel to *DK* (light solid line) to determine point of intersection *V* with segment *CP* (see the inset in Figure 5.31).

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By Apollonius III.48 (see Section 3.2. The Ellipse), the normal to *DK* from point *F* to point *P* (let PF = h) is he bisector of angle *SPH*. That is, triangle *IPH* is an isosceles triangle and so (by Euclid I.6) IP = PH. Since SC = CH, then by Thales' Intercept Theorem (Theorem 1.1) we have SE = EI. By Apollonius III.52 (the alternative definition of an ellipse from Section 3.2. The Ellipse), we have SE + EI + IP + PH = 2a. Substituting *El* for *SP* and substituting *IP* for *PH*, we have: EP = EI + IP = a. (5.55)





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Newton's Lemma, continued 3

Proof (continued).



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If the ellipse were in fact a circle, then by Euclid III.35 (see Section 2.2. Book III) we would have $GV \cdot VP = QV^2$ (here we would need to extend QV until it is a chord of the circle; see the inset in Figure 5.31).



We now take the ellipse (along with the segments GV, VP, and QV) and stretch/shrink it in the direction of the blue "diameter" by a factor 1/c, and stretch/contract it in the direction of the red "diameter" by a factor of 1/d. This results in a circle of radius 1. Since GV and VP lie along the blue diameter then these are transformed to the circle with lengths GV/cand VP/c. Since QV lies on a line parallel to the blue diameter then this is transformed to the circle with length QV/d. So from the above equation $GV \cdot VP = QV^2$ for the chords of a circle, we have

$$\frac{GV \cdot VP}{c^2} = \frac{QV^2}{d^2}, \text{ or: (3) } VP = \frac{c^2}{GV} \cdot \frac{QV^2}{d^2}.$$



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Next, we express *VP* in terms of *RQ*, and express *QV* in terms of *QT* (see the inset). Now triangle *XVP* is similar to triangle *ECP*, because they share an angle at point *P* and angles *PXV* and *PEC* are corresponding angles for parallel line segments *QV* and *EF* with transversal *CP* (see Figure 1.7 center in Section 1.3. Properties of Angles); so the three angles of triangles *XVP* and *ECP* are the same. So XP/VP = EP/CP or (since EP = a by (5.55)) XP/VP = a/c, or: (2) $XP = VP \cdot a/c$.



Fig. 5.31. Newton's proof of his lemma

Triangle QTX is similar to triangle *PFE* since both of these are right triangles and angles QXT and *PEF* are alternate interior angles (or "parallel angles," see Figure 1.7 let in Section 1.3. Properties of Angles) for parallel line segments QX and *EF* with transversal *EX*; so the three angles of triangles QTX and *PFE* are the same. Also QX/QT = PE/PF or (since PF = h) QX/QT = a/h, or:

(6)
$$QX = QT \cdot a/h$$
.



Fig. 5.31. Newton's proof of his lemma

We now consider PQ "infinitely small." In contemporary (rigorous) terms, we consider a limit as $Q \rightarrow P$ and Q is on the ellipse. We see from the inset of Figure 5.31 that:

(1)
$$RQ \approx XP$$
, (4) $GV \approx GP = 2c$, (5) $QV \approx QX$.

In fact, each of these can be written as equalities in the limit.

Proof (continued). We now have:

$$RQ \approx XP \text{ by } (1)$$

$$= VP \cdot \frac{a}{c} \text{ by } (2)$$

$$= \left(\frac{c^2}{GV} \cdot \frac{QV^2}{d^2}\right) \cdot \frac{a}{c} \text{ since by } (3) VP = \frac{c^2}{GV} \cdot \frac{QV^2}{d^2}$$

$$\approx \frac{c^2}{2c} \cdot \frac{QV^2}{d^2} \cdot \frac{a}{c} \text{ since by } (4) GV \approx GP = 2c$$

$$\approx \frac{c^2}{2c} \cdot \frac{QX^2}{d^2} \cdot \frac{a}{c} \text{ since by } (5) QV \approx QX$$

$$= \frac{c^2}{2c} \cdot \frac{(QT \cdot a/h)^2}{d^2} \cdot \frac{a}{c} \text{ since by } (6) QX = QT \cdot a/h$$

$$= \frac{a^3}{2h^2d^2} \cdot QT^2.$$

Proof (continued). We now apply Apollonius VII.31, which is stated in Exercise 3.5.2 as: "All parallelograms circumscribed about any conjugate diameters of a given ellipse are equal."



Notice that 1/4 of the areas of the inscribed rectangles above are *hd* (left) and *ab* (right). Since the areas of the parallelograms are the same by Apollonius VII.31, then hd = ab.

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Proof (continued). So with hd = ab, we have from the approximation $RQ \approx \frac{a^3}{2h^2d^2} \cdot QT^2$ (established above) that

$$RQ \approx rac{a^3}{2h^2d^2} \cdot QT^2 = rac{a^3}{2a^2b^2} \cdot QT^2 = rac{a}{2b^2} \cdot QT^2.$$

That is, $RQ \approx (\text{Constant}) \cdot QT^2$ where $\text{Constant} = \frac{a}{2b^2}$ and the constant is independent of the position of *P* on the ellipse, as claimed.

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Proof. By equation (5.53) we have that the force f is proportional to RQ. By Newton's Lemma, RQ is approximately proportional to QT^2 (and so f is approximately proportional to QT^2). The area of the triangle SPQ in Figure 5.30 right is $SP \cdot QT/2$ and this area is a constant for a fixed Δt , as seen in the proof of Kepler's Second Law (Theorem 5.8).

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Proof (continued). Since $SP \cdot QT/2$ is constant, then QT is inversely proportional to SP. Hence f is approximately inversely proportional to SP = r. That is, $f \approx \frac{Constant}{r^2}$. By taking a limit as $\Delta t \rightarrow 0$, we have $Q \rightarrow P$ and the approximation becomes precise.